# HP-41 Module: Areas, Series & Sums. Fractional Integro-Differentiation



## Overview

This module includes a selection of functions and FOCAL routines mainly focused on Series and Sums field and other related subjects. For the most part the routines are taken from Jean-Marc Baillard extensive web site, although some others are takes from Poul Kaarup's collection as well. A few are already available in the SandMath Module –even if this version is a more portable implementation that suits itself better for Clonix/NoVRAM owners.

The initial section of the module covers the simple sums of integers and integer powers. This is followed by simple explicit sums for single, double, triple and multiple series; recursive term sums and Euler transformations. Examples are also provided in the FAT for quick familiarization. The second section includes a set of MCODE functions and FOCAL routines dealing with area calculations for several geometric figures, such as circles and triangles, as well as areas and diagonal lengths of cyclic and regular polygons.

The last section takes a relative large part of the FAT and deals with Fractional Integro-differentiation of a few Elemental and Special functions. This technique is based on the Hyper-Geometric function and consolidates an impressive set of capabilities under the same approach.Without further ado, see below the list of functions included in the module:

XROM	Function	Description	Input	Author
18,00	-SERIES 1A	Rounded Comparison	x,y in X,Y	Ángel Martin
18,01	СНЅҮХ	Sign Change of Y by X: Y*(-1)^X	Value in Y, n in X	Ángel Martin
18,02	NCR	Combinations of N in sets of R	N in Y, R in X	Ángel Martin
18,03	"#"	Auxiliary for SOLVE	Under program control	Ángel Martin
18,04	Σ0	Sum of mantissa digits	argument, X	Ángel Martin
18,05	Σ1	Sum of first N integers	argument, X	Poul Kaarup
18,06	Σ1/Ν	Harmonic Numbers	argument, X	Ángel Martin
18,07	Σ2	Sum of Squares of Numbers	argument, X	Poul Kaarup
18,08	Σ3	Sum of Cubes of Numbers	argument, X	Poul Kaarup
18,09	ΣΝ^Χ	Generalized Faulhaber's Sum	N in Y, exponent In X	Ángel Martin
18,10	"ΣυΜ″	Single Series Sum (Explicit)	arguments in X, Y, ALPHA	JM Baillard
18,11	"ΣΣυΜ"	Double Series Sum (Explicit)	argument in X, ALPHA	JM Baillard
18,12	"ΣΣΣυΜ"	Triple Series Sum (Explicit)	arguments in Y,X, ALPHA	JM Baillard
18,13	"ΣυΜΕ	Euler Transformation	argument in X, Y, Z, ALPHA	JM Baillard
180,14	"ΣUMR"	Single Series Sum (Iterative)	arguments in X, Y, ALPHA	JM Baillard
18,15	"ΝΣυΜ0"	Multiple Series Sum (Explicit)	bbb.eee in X, data in registers	JM Baillard
18,16	"NT"	Example for N $\Sigma$ UM0	n/a	Martin-Baillard
18,17	"Т"	Example for $\Sigma$ UM	n/a	Martin-Baillard
18,18	"TE"	Example for $\Sigma$ UME	n/a	Martin-Baillard

18,19	"TR"	Example for $\Sigma$ UMR	n/a	Martin-Baillard
18,20	"TT"	Example for $\Sigma\Sigma UM$	n/a	Martin-Baillard
18,21	"TTT"	Example for $\Sigma\Sigma\Sigma$ UM	n/a	Martin-Baillard
18,22	-AREAS_1B	ALPHA Integer part	Argument in X	Fritz Ferwerda
18.23	BRHM	Brhamagupta formula	4 Sides in stack	Ángel Martin
18.24	CIRCLE	Circle through three points	Coordinates in R01-R06	Ángel Martin
18.25	DIAG	Diagonal formula	Used in CPLD	JM Baillard
18.26	HERON	Heron formula	Triangle sides in Z, Y, Z	Ángel Martin
18.27	RPG1	Regular Polygon from sides	# sides in Y, length in X	Poul Kaarup
18.28	RPG2	Regular polygon from circle	# sides in Y, radius in X	Poul Kaarup
18.29	"3PNTS"	Driver for CIRCLE/Areas	Prompts for coordinates	Ángel Martin
18.30	"CCPA"	Complex Cyclic Polygon Area	Parameters in stack and R00	JM Baillard
18.31	"CCPA+"	Driver for CCPA	Uses Newton Method	Ángel Martin
18.32	"CPLA"	Complex Cyclic Polygon Area	With all Sides known	JM Baillard
18.33	"CPLA+"	Driver for CPLA	Uses SOLVE	Ángel Martin
18.34	"CPLD"	Complex Cyclic Polyg. Diagonals	bbb.eee in X, data in Registers	JM Baillard
18.35	"CPLD+	Driver for CPLD	Prompts for data entry	Ángel Martin
18.36	"STLA"	Star Polygon Area	Parameters in Stack	JM Baillard
18.37	"STLA+"	Driver for STLA	Prompts for data entry	Ángel Martin
18.38	-ITG/DIFF	Generalized Hypergeometric	Data in Registers and Stack	JM Baillard
18.39	1/GM	Reciprocal of Gamma	argument in X	JM Baillard
18.40	PSI	Digamma Function	argument n X	JM Baillard
18.41	"DAIRY"	Airy Function Intg-diffm	μ in Y, x in X	JM Baillard
18.42	"DEI"	Exponential Intg. Intg-Diff.	$\mu$ in Y, x in X	JM Baillard
18.43	"DERF"	Error Function IntgDiff	μ in Y, x in X	JM Baillard
18.44	"DCHI"	Hyp Cos Integral Intg-Diff.	μ in Y, x in X	JM Baillard
18.45	"DCH"	Hyp. Cosine Intg-Diff.	μ in Y, x in X	JM Baillard
18.46	"DCI"	Cosine Integral Intg-Diff.	μ in Y, x in X	JM Baillard
18.47	"DCOS"	Cosine Intg-Diff	μ in Y, x in X	JM Baillard
18.48	"DCX"	Fresnel Cosine Integral Intg-Diff.	μ in Y, x in X	JM Baillard
18.49	"DEXP"	Exponential Intg-Diff.	μ in Y, x in X	JM Baillard
18.50	"DHMT"	Hermite Function Intg-Diff.	$\mu$ in Z, n in Y, x in X	JM Baillard
18.51	"DINX"	Bessel I function Intg-Diff.	$\mu$ in Z, n in Y, x in X	JM Baillard
18.51	"DJNX"	Bessel J function Intg-Diff.	$\mu$ in Z, n in Y, x in X	JM Baillard
18.53	"DKNX"	Mod. Bessel K function Intg-Diff.	$\mu$ in Z, n in Y, x in X	JM Baillard
18.54	"DKUM"	Kummer function intg-Diff.	$\mu$ in Y, x in X, a in R00, b in R01	JM Baillard
18.55	"DLANX"	Lagrange Function Intg-Diff.	$\mu$ in T, a in Z, n in Y, x in X	JM Baillard
18.56	"DLN"	Natural Log Intg-Fiff.	μ in Y, x in X	JM Baillard
18.57	"DSB1"	Spherical Bessel Funct. 1 <sup>st</sup> . Kind	μ in Z, n in Y, x in X	JM Baillard
18.58	"DSHI"	Hyp. Sine Integral inth-Diff.	μ in Y, x in X	JM Baillard
18.59	"DSH"	Hyperbolic Sine Intg.Diff.	μ in Y, x in X	JM Baillard
18.60	"DSI"	Sine integral Intg-Diff.	μ in Y, x in X	JM Baillard
18.61	"DSIN"	Sine Intg-Diff	μ in Y, x in X	JM Baillard
18.62	"DSX"	Fresnel Sine Integral Intg-Diff.	μ in Y, x in X	JM Baillard
18.63	"DYNX"	Mod. Bessel Y function Intg-Diff.	μ in Z, n in Y, x in X	JM Baillard

### **1**–Sums and Series

The first section includes several functions to calculate sums of integer powers, as well as simple methods to sum series given their general term in explicit or recurrent form.

• **∑0** is a small divertimento useful in pseudo-random numbers generation. It simply returns the sum of the mantissa digits of the argument – at light-blazing speed using just a few MCODE instructions. More about random numbers will be covered in the Probability/Stats section later on.

Example: calculate the sum of all digits of the HP-41's rendition of pi:

PI, XEQ "**∑0**" => 40.00000000

•  $\Sigma 1/N$  calculates the Harmonic number of the argument in X, that is the sum of the reciprocals of the <u>natural</u> numbers (which excludes zero) lower and equal to n. It will be used in the calculation of the Kelvin functions and the Bessel functions of the second kind, K(n,x) and Y(n,x).

$$H_n = \sum_{k=1}^n \frac{1}{k}.$$

Example: calculate H(5) and H(25).

5, XEQ "Σ1/N"	=> 2.283333333
25, XEQ "Σ1/N″	=> 3.815958178

• Σ1, Σ2, Σ3 are implemented to calculate the sum of integer powers directly based on the corresponding formulas. The number of terms to sum is expected to be in the X- register. Functions calculate the linear sum using the triangular formula; the sum of squares using the pyramidal formulas; and the sum of cubes also using the pyramidal formulas.

$$T_n = \sum_{k=1}^n k = 1 + 2 + 3 + \dots + n = \frac{n(n+1)}{2} = \binom{n+1}{2}$$
$$P_n = \sum_{k=1}^n k^2 = \frac{n(n+1)(2n+1)}{6} = \frac{2n^3 + 3n^2 + n}{6}.$$

Example: Calculate the sum of the first 10 natural numbers and their squares and cubes:

10, **∑1** quickly returns: 55.0000000 LASTX, **∑2**″ => 385.000000 LASTX, **∑3** => 3,025.000000

•  $\Sigma N^X$  Calculates a generalized value of the Faulhaber's formula for integer values of x. – The few first integer values of x have explicit formulas for the result (which are used in the functions described above), but that's not the case for a general value - which can also be non-integer. Obviously for x=-1 this function returns identical results than  $\Sigma 1/N$ , albeit slower due to the additional complexity of the definition of the term.

<u>Example:</u> Check the triangular (x=1) and pyramidal (x=2) formulas for n=10 – which are particular cases of the Faulhaber's Formula, involving Binomial coefficients and Bernoulli's numbers. See the link below for details: <u>http://en.wikipedia.org/wiki/Faulhaber%27s\_formula</u>

10, ENTER^, 1, XEQ "ΣN^X"	=> 55.0000000
10, ENTER^, 2, XEQ "∑N^X"	=> 385.000000

And using the convention B(1) = 0.5 the formula is:

$$S_m(n) = \frac{1}{m+1} \sum_{k=0}^m \binom{m+1}{k} B_k \ n^{m+1-k},$$

Which could be programmed using a few of the SandMath functions, albeit it would be considerably slower due to the impact of the Zeta algorithms (part of Bernoulli's) – kicking in for n>4.

• **CHSYX** is related to the same subject, and in general relevant to the summation of alternating series – It can be regarded as an extension of **CHS** but dependent of the number in X. Its expression is:

CHS(y,x)=  $y^{*}(-1)^{x}$ , and thus changing the sign of Y when the number in X is odd.

• **NPR** calculates Permutations, defined as the number of possible different arrangements of N different items taken in quantities of R items at a time. No item occurs more than once in an arrangement, and different orders of the same R items in an arrangement are counted separately. The formula is:

$$\frac{n!}{(n-k)!}$$

This last two functions will be used as subroutines in the series sums programs described next.



#### Sums of Series using their General Terms

The following programs allow you to obtain the value of simple, double and triple series for which the general term is known, either as explicit expression or as a recurrent form.

#### Simple Series.

 $\Sigma UM$  calculates a single series sum, that is:

 $S = (u_k + u_{k+1} + u_{k+2} + u_{k+3} + \dots) = \Sigma u_n$  for  $n \ge k$ 

defined by the general term  $u_n = f(n)$  where f is a known (i,e, explicit) function.

A program which computes Un = f(n) is required as a subroutine. It is done assuming n is in X-register upon entry. The module includes the routine "**T**" as example for this case, with Un=1/n!

ALPHA, "TR", 1, ALPHA, XEQ "∑UM" -> X = 1.718281830 = R01 (in 9 seconds)

**\SigmaUMR** calculates the same sum when Un is given as a recurrence expression with an initial known value, that is:  $u_k$  is given and  $u_{n+1} = f(u_n;n)$  where f is a known function.

a program which computes un+1 = f(Un; n) is required as a subroutine. It is done assuming Un is in X-register and n is in Y-register upon entry. The module includes the routine "**TR**" as example for this case, with  $u_1 = 1$  and  $u_{n+1} = u_n/(n+1)$ 

ALPHA, "TR", ALPHA, 1, ENTER^, XEQ " $\Sigma$ UMR" -> X = 1.718281830 = R01 (in 7 seconds)

Note that on both cases the initial index can also be zero, assuming that's compatible with the definition of Un, which adds more flexibility to the routine. In both cases the function needs to be programmed under a global label, and *its name is expected to be in the ALPHA register when the routines are called.* 

STACK	INPUTS	OUTPUTS	STACK	INPUTS	OUTPUTS
Y	/	ΣUΜ	Y	k	ΣUMR
Х	k	ΣUΜ	X	u <sub>k</sub>	ΣUMR
ALPHA	F.NAME	F.NAME	ALPHA	F.NAME	F.NAME

#### Euler Transformation.

**SUME** calculates the same sum making use of the *Euler Transformation* to accelerate the convergence of alternating series:  $S = u_0 - u_1 + u_2 - u_3 + \dots + (-1)^n u_n + \dots$ 

The sum is re-written in function of the binomial coefficients,  $C_n^p = n! / (p! (n-p)!)$  as follows:

$$S = a_0/2 + (C_1{}^1 a_0 - C_1{}^0 a_1)/2^2 + (C_2{}^2 a_0 - C_2{}^1 a_1 + C_2{}^0 a_2)/2^3 + (C_3{}^3 a_0 - C_3{}^2 a_1 + C_3{}^1 a_2 - C_3{}^0 a_3)/2^4 + \dots$$

This may produce superb acceleration but it can also fail.

A program which computes Un = |f(n)| is required as a subroutine, but *without the alternating sign.* It is done assuming n is in X-register upon entry. The module includes the routine "**TE**" as example for this case, with  $f(n) = (n+1)^{-1/2}$ 

ALPHA, "TE", ALPHA, XEQ "∑UME" -> X = 0.6048986431 = R01

For this particular example the error is -3 10-10 (!), and only 28 terms are calculated (taking about 6.5 minutes to converge). Without an acceleration method, more than 1,018 terms would be necessary to achieve the same accuracy... which execution time would be much greater than the age of the Universe.

STACK	INPUTS	OUTPUTS
Х	/	ΣυΜΕ
ALPHA	F.NAME	F.NAME

#### Double and Triple Series.

**Solution** calculates a <u>double series sum</u>, that is:  $S = \Sigma\Sigma (u_{n;m})$ ; for  $n \ge n_0$ ;  $m \ge m_0$ 

As before, the general term needs to be programmed under a separate subroutine using a global label. This assumes that "n" is in the X-register and "m" in the Y-register upon entry. The two initial indices are expected to be in the stack and the label name must be in the ALPHA register as well.

STACK	INPUTS	OUTPUTS
Y	m <sub>0</sub>	ΣΣ
Х	<b>n</b> 0	ΣΣ
ALPHA	F.NAME	F.NAME

The program uses data registers R00 to R05, which therefore should not be used in the definition of the general term. The module includes the routine "**TT**" as example for this case, with the expression:  $f(n;m) = 1 / (n^n m!)$ 

ALPHA, "TT", ALPHA, 1, ENTER^, XEQ "ΣΣUM" -> X = 2.218793264 = R03 (in 1 min 18s)

 $\Sigma\Sigma\Sigma$  ( $u_{n;m;p}$ ); for  $n \ge n_0$ ;  $m \ge m_0$ ;  $p \ge p_0$ 

As before, the general term needs to be programmed under a separate subroutine using a global label. This assumes that "n" is in the X-register "m" in the Y-register, and p is in the Z-register upon entry. The three initial indices are expected to be in the stack and the label name must be in the ALPHA register as well.

STACK	INPUTS	OUTPUTS
Z	$p_0$	/
Y	$m_0$	ΣΣΣ
Х	n <sub>0</sub>	ΣΣΣ
ALPHA	F.NAME	F.NAME

The program uses data registers R00 to R07, which therefore should not be used in the definition of the general term. The module includes the routine "**TTT**" as example for this case, with the expression:  $f(n;m;p) = 1 / (n^n m! (p!)^2)$ 

ALPHA, "TTT", ALPHA, 1, ENTER^, ENTER^, XEQ "ΣΣΣUM" -> X = 2.839135243 = R04 (6 min 5s)

With an error  $\varepsilon = -7 \text{ E-9}$ 

#### Multiple Series

**N** $\Sigma$ **UMO** calculates a <u>multiple series sum</u>, with k internal summations that all start in zero - that is:

**S** =  $\Sigma\Sigma...\Sigma$  [ **U**(n<sub>1</sub>;n<sub>2</sub>; ....; n<sub>k</sub>) ] with n<sub>1</sub>>= 0; n<sub>2</sub>>= 0; .....; n<sub>k</sub>>= 0

Obviously now the number of indices will be in the X register, and there are no initial indices – which are assumed to be zero. This may need you to re-write the expression of the general term to make it compatible with this condition.

This is the most limiting requirement for this program, which is not suitable for cases that have mutual dependencies between the initial indexes.

Here too, the general term needs to be programmed under a separate subroutine using a global label, which needs to be entered in ALPHA. This assumes on entry that "n1" is in the R01 register, "n2" in the R02 register, "n3" in R03, and successively so until completing the number of variables.

STACK	INPUTS	OUTPUTS
Х	k	ΝΣΟ
ALPHA	F.NAME	/

Only the synthetic registers {M,N,O} are used by the program. The module includes the routine "**NT**" as example for this case, with the expression used in the triple aeample :  $f(n;m;p) = 1 / (n^n m! (p!)^2)$ .

We therefore need to change it to start at the zero indexes for the three variables, i.e. must make a change of arguments to reduce to the standard:  $n \ge 0$ ;  $m \ge 0$ ;  $p \ge 0$  by replacing n with (n+1); m with (m+1); p with (p+1):  $f(n;m;p) = 1 / \{(n+1)^{(n+1)} (m+1)! [(p+1)!]^2\}$ 

ALPHA, "NT", ALPHA, 3, XEQ "NΣUM0" -> X = 2.839135243 = R04 (8 min 39 s)

$$F(x,v) = \sum_{k=0}^{n} \sum_{m=0}^{k} \sum_{m=0}^{k} \alpha_{m1,m2,k} R^{k} y^{k-m1} (1-y)^{m1} z^{k-m2} (1-z)^{m2}$$

### 2 – Areas of Polygons

The second section of the module includes several MCODE functions for triangles, cyclic quadrilaterals and even non-regular polygons.

• **CIRCL** calculates the radius of a circle passing thru three data points, using the point x,y coordinates. The values are expected to be stored in R01-R07. Besides that, it'll also return in the Y-register the area of the circumscribed triangle defined by the three points.

Example: Calculate the radius of the circle passing thru P(5,1), Q(6,2), and R(5,3)



The results are:

XEQ "CIRCL" => r=1,00000000, X<>Y => A=1,00000000

The input sequence starts with the abscissa of P1 in R01.

Note that you can use the routines **IN** and **INPUT** available in the SandMath to populate the registers automatically.

• **HERON** calculates the area of a triangle knowing its three sides, using Heron's formula. Just enter the sides values in the stack, and execute the function. The result is stored in X, with the original side saved in LastX. The rest of the stack is unchanged.

Let the triangle ABC with 3 known sides { a , b , c } and s = (a+b+c)/2 the semi-perimeter

Heron's formula is: Area = [s(s-a)(s-b)(s-c)]1/2

<u>Example:</u> a = 2, b = 3, c = 4 Type: 2, ENTER<sup>^</sup>, 3, ENTER<sup>^</sup>, 4, XEQ "HERON" => Area = 2.904737510

Note: the function **CIRCL** described above makes use of the HERON formula internally after it first calculates the triangle sides from the point coordinates.

• **BRHM** is related to it, but the calculation for the area of the cyclic quadrilateral - using Brhamagupta's formula. Just enter the four values in the stack and execute the function. The result is stored in X, with the original side saved in LastX. The rest of the stack is unchanged.

Let a, b, c, and d be its sides lengths, and the semi-perimeter s = (a + b + c + d)/2. The area A of the cyclic quadrilaterals:

A = [(s-a).(s-b).(s-c).(s-d).]1/2

*Example:* a = 4 , b = 5 , c = 6 , d = 7

Type: 4, ENTER<sup>^</sup>, 5, ENTER<sup>^</sup>, 6, ENTER<sup>^</sup>, 7, XEQ "BRHM" => Area = 28.98275349



• **PG1** Calculates the area of a regular polygon when its side length is known. Input parameters are the number of sides in Y, and the side length in X. The result is left in the X register.

Example: calculate the area of a triangle with side length a=5 m

3, ENTER^, 5, XEQ "PG1" -> 10.83 m^2

• **PG2** Performs the same calculation but using the radius of the circumscribed circle instead of the side length. Same order of parameters for input, with the number of sides in Y.

Example: calculate the area of a triangle circumscribed in a circle with radius r=5 m

3, ENTER^, 5, XEQ "PG2" -> 32.48 m^2

Finally, the Routine "**3PNTS**" is a FOCAL driver for functions **CIRCLE**. You can use it to enter the coordinates of the three points into data registers R01-R06, presented as three screens with the prompts:



Once this is accomplished the program offers you a choice for the value to calculate next, either the triangle area or the circle radius. You can Also press [E] to start over with a new set of three points.



If used on the example listed above, it returns the following results:



#### Convex Cyclic Polygons.

And what about non-regular polygons, you may wonder? Well, those are the subject of the following set of FOCAL routines about to be described.

Programs **"CCPA"** and **"CPLA"** compute the area A and the circumradius R of a convex cyclic polygon assuming all the sides lengths are known. Moreover, we also assume that the *center* of the circumcircle is *inside* the polygon.

If  $\{a_1, a_2, \dots, a_n\}$  are the sides lengths and  $\{\mu_1, \mu_2, \dots, \mu_n\}$  are the corresponding central angles, we have to solve the system of (n+1) equations:

2.R sin  $\mu_1/2 = a_1$ 2.R sin  $\mu_2/2 = a_2$ 2.R sin  $\mu_n/2 = a_n$  $\mu_1 + \mu_2 + ... + \mu_n = 360^\circ$ 

Performing a few substitutions lead to an expression with the radius as single unknown, to be resolved iteratively using any root-finding method:

 $asin (a_1/2R) + asin (a_2/2R) + \dots + asin (a_n/2R) = 180^{\circ}$ 

After finding R, the Area is given by :  $A = (R^2/2) (Sin a_1 + Sin a_2 + ... + Sin a_n)$ 

There are two versions included in the module – "**CPLA**" uses the SOLVE function in the Advantage and "**CCPA**" uses a built-in root finder based on Newton's method. Each one has advantages and shortcomings, as usual.

#### Drivers for Data Entry.

The routines expect the sides of the polynomial already stored in contiguous data registers, and the control word "bbb.eee" in the X register before you call the routine. For your convenience, a driver routine is also included that prompts for the side values and does the storing for you, Using "**CPLA+**" or "**CCPA+**", all you need to do is enter the values at each prompt, and once completed it'll direct the execution to the corresponding data engine downstream.

Example. Find the area of a convex cyclic polygon with sides: 4, 5, 6, 7, 8, 9, 10

XEQ "CPLA+"	"N=?"
7, R/S	"d1=?"
4, R/S	"d2=?"
5, R/S	"d3=?"
6, R/S	"d4=?"
7, R/S	"d5=?"
8, R/S	"d6=?"
9, R/S	"d7=?"
10, R/S	-> 174.6757940 the area, and
X<>Y	-> 8.143816980 the radius

#### Diagonal Lengths.

**CPLD** calculates the diagonal lengths of a convex cyclic polygon with its side lengths known. The method used involves solving a linear system of (n-3) equations with (n-3) unknowns, which is solved by successive approximations. The convergence is linear only – which results in relative longer execution times.

The iteration starts with all diagonals lengths = 0, which is very simplistic. The successive sums of the differences between 2 consecutive approximations (in absolute values) are displayed. They should tend to zero, however, the termination criterion may lead to an infinite loop. Since there are 319 registers at most, "CPLD" can find the diagonals lengths of a 159-gon – but the execution time will not be small without an emulator.

**CPLD** expects the number of sides in R00, and the sides of the polynomial already stored in contiguous data registers starting with R01 until Rn+1. Then you must provide the control words "bbb.eee" indicating the location of the data registers that store the side lengths.

STACK	INPUT	OUTPUTS
Х	/	bbb.eee

For your convenience, a driver routine is also included that prompts for the side values and does the storing for you, Using "**CPLD+**" all you need to do is enter the values at each prompt, and once completed it'll direct the execution to CPLD downstream.

Example: Find the diagonals lengths of a convex cyclic hexagon with sides: 4, 5, 6, 7, 8, 9

XEQ "CPLD+"	"N=?"
6, R/S	"a1=?"
4, R/S	"a2=?"
5, R/S	"a3=?"
6, R/S	"a4=?"
7, R/S	"a5=?"
8, R/S	"a6=?"
9, R/S	shows estimations> convergence
	d8=8.46278437
R/S	d9=12.12358502
R/S	d10=12.97690535

There are in fact n(n-3)/2 diagonals whose lengths may be obtained by "rotating" the sides lengths in registers R01 to R06 and we have similarly: n(n-3)/2 = 9 if n = 6

$d_4 = 9.998827970$	d <sub>7</sub> = 11.30861231
$d_5 = 13.01214483$	$d_8 = 13.06010803$
$d_6 = 11.49035918$	$d_9 = 12.32872367$

Finally, the MCODE function **DIAG** is used internally by **CPLD** to speed-up the calculations. It computes the following expression, assuming x, y, z, t are in registers X, Y, Z, T upon entry

SQRT [ { x.y ( $z^2 + t^2$ ) + z.t ( $x^2 + y^2$ ) } / (x.y + z.t) ]

#### Regular Star Polygons

A star polygon  $\{n/k\}$ , with n,k positive integers, is a figure formed by connecting with straight lines every "k-th" point out of n regularly spaced points lying on a circumference. The number k is called the polygon density of the star polygon. Without loss of generality, take k<n/2. The star polygons were first systematically studied by Thomas Bradwardine.



If k=1, a regular polygon {n} is obtained. Special cases of {n/k} include {5/2} (the pentagram), {6/2} (the hexagram, or star of David), {8/2} (the star of Lakshmi), {8/3} (the octagram), {10/3} (the decagram), and {12/5} (the dodecagram).

"STLA" computes the area A, the perimeter P, the inradius r and the circumradius R of a regular star polygon { n/k } from its edge length a

Formulae:

 $\begin{array}{l} A = n \ R^2 \ Sin \ (180^\circ/n) \ Cos \ (180^\circ \ k/n) \ / \ Cos \ [180^\circ(k-1)/n] \\ a = 2.R \ Sin \ (180^\circ \ k/n) \\ r \ = \ R \ Cos \ (180^\circ \ k/n) \\ P \ = 2.A \ / \ r \end{array}$ 

The table on the left shows the input and output Required by the program – easy does it!

STACK	INPUTS	OUTPUTS
Т	/	R
Z	а	r
Y	n	Р
X	k < n/2	А

#### Examples:

• $a = 1$ , $n = 5$ , $k = 2$			
1 ENTER^, 5 ENTER <sup>4</sup>	^, 2 XEQ " <b>STLA</b> "	->	A = 0.310270701
	RDN		P = 3.819660113
	RDN		r = 0.162459848
	RDN		R = 0.525731112
• $a = 1$ , $n = 10$ , $k = 3$			
1 ENTER^. 10 ENTER	. 3. R/S	->	A = 0.857567126
1 200200 20020	RDN	-	P = 4 721359547
	RDN		r = 0.363271264
	RDN		R = 0.618033989
	RDN		K = 0.010055505
• $a = \pi$ , $n = 41$ , $k = 13$			
PI ENTER^, 41, ENTER	R^, 13, R/S	->	A = 9.855571194
	RDN		P = 19.37713086
	RDN		r = 1.017237409
	RDN		R = 1.871409374

This program works in all angular modes, however, DEG mode should be preferable.

If k = 1, we get the convex regular n-gon. For instance, with a = 1, n = 5, k = 1, "STLA" returns what corresponds to the regular pentagon.

A = 1.720477401 P = 5 r = 0.688190960R = 0.850650808

#### Driver Program.

Here too you have a convenient driver program to guide you thru the data entry process: program "STLA+" will prompt for the input values and will present the results sequentially after the calculations are done.



### **3**–Fractional Integro-Differentiation

If  $\boldsymbol{\mu}$  is a real number (integer or fractional) and a function f is defined by a power series:

 $f(x) = \sum_{k=0,1,2,..} \{ c_k x^k \}$ 

then its fractionalintegro-differentiation may be computed by:

 $d^{\mu} f / dx^{\mu} = D^{\mu} f(x) = \sum_{k=0,1,2,\dots} \{ c_k [ \Gamma(k+1) / \Gamma(k+1-\mu) ] x^{k-\mu} \}$ 

If the function may be expressed in terms of hypergeometric  ${\sf functions}_p F_q$ , the following relation is very useful too:

 $D^{\mu}{}_{p}F_{q}\;(\;a_{1},\;\ldots\;,\;a_{p}\;;\;b_{1},\;\ldots\;b_{q}\;;\;x\;)=x^{-\mu}\Gamma(b_{1})....\Gamma(b_{q})_{p+1}F^{\sim}{}_{q+1}\;(1,\;a_{1},\;\ldots\;,\;a_{p}\;;\;1-\mu,\;b_{1},\;\ldots\;b_{q}\;;\;x)$ 

where  $\Gamma$  is Euler's Gamma function, and  ${}_{p}F^{\sim}{}_{q}$  is the regularized generalized hypergeometric function. This function was already included in the SandMath module, and it is now copied here for convenience – tucked away under one of the section headers so that a dedicated FAT entry was not required.

As a trivial corollary,  $D^0 f(x) = f(x)$ . Positive integer numbers (i.e. natural numbers) represent the successive derivatives of the function. Also if  $\mu$ = -1, -2, -3, .... the results are the repeated integrals of the function f - usually those that vanish for x = 0.

#### Formulae and details.

Besides the special functions in the module there are several elementary functions included as well. Note that all program names begin with the 'D" letter, followed by the common acronym that designates the function's name.

As usual when a function is evaluated by a power series, the results are not very accurate for large arguments; they may even be meaningless ... unless all the terms have the same sign!

For data entry,  $\mu$  is always to be entered first, then the order/index - if any - and finally, x in register X: (\*)

Stack	Input#1	Input#2	Input#3	Result
Т	/	/	μ	/
Z	/	μ	а	/
Y	μ	n	n	/
X	X	X	X	$(D^{\mu} f)(x)$

(\*) as single exception, for Kummer's functions R01 and R02 must also be initialized first with the function parameters a & b.

You're encouraged to visit JM's original page located at:<u>http://hp41programs.yolasite.com/fracintegrdiff.php</u>

### A few elementary functions.

•	Hyperbolic Sine	D <sup><math>\mu</math></sup> Sinh x = 2 <sup><math>\mu</math>-1</sup> sqrt( $\pi$ ) x <sup>1-<math>\mu</math></sup> <sub>1</sub> F <sup><math>\sim</math></sup> <sub>2</sub> (1; (2- $\mu$ )/2, (3- $\mu$ )/2; x <sup>2</sup> /4)
•	Hyperbolic Cosine	$D^{\mu} \operatorname{Cosh} x = (2/x)^{\mu} \operatorname{sqrt}(\pi) {}_{1}F^{2} (1; (1-\mu)/2, (2-\mu)/2; x^{2}/4)$
•	Sine	$D^{\mu}$ Sin x = 2^{\mu - 1} sqrt(\pi) x^{1 - \mu} {}_1F^{\sim}_2 ( 1 ; (2- $\mu$ )/2 , (3- $\mu$ )/2 ; -x²/4 )
•	Cosine	$D^{\mu} \cos x = (2/x)^{\mu} \operatorname{sqrt}(\pi) {}_{1}F^{2}(1;(1-\mu)/2,(2-\mu)/2;-x^{2}/4)$
•	Exponential	$D^{\mu} Exp \ x = x^{-\mu}{}_{1}F^{-}{}_{1} (1; 1-\mu; x)$
•	<u>Logarithm</u>	$D^{\mu} Ln x = x^{-\mu} FC^{(\mu)}_{log}(x)$
	where $FC^{(\mu)}_{log}$ and $FC^{(\mu)}_{log}$	$g(x) = (-1)^{\mu-1} (\mu-1) !  \text{if } \mu \text{ is a positive integer,}$ $g(x) = [\operatorname{Ln} x - \operatorname{Psi}(1-\mu) - \gamma] / \Gamma(1-\mu)  \text{otherwise }.$

Psi = Digamma Function;  $\gamma$ = Euler's constant =0.5772156649... and  $\Gamma$ =Gamma Function.

#### Examples:

Hyperbolic Sine:     3.14 ENTER^, 1.28	XEQ "DSH"	-> $D^{3.14}$ Sinh ( 1.28 ) = 1.999005451
Hyperbolic Cosine:     3.14 ENTER <sup>^</sup> , 1.28	XEQ "DCH"	-> $D^{3.14}$ Cosh ( 1.28 ) = 1.502958219
• Sine: 3.14 ENTER^, 1.28	XEQ "DSIN"	-> $D^{3.14}$ Sin (1.28) = -0.019142092
• Cosine: 3.14 ENTER^, 1.28	XEQ "DCOS"	-> $D^{3.14}$ Cos (1.28) = 0.888787267
<ul> <li>Logarithm:</li> <li>3.14 ENTER<sup>^</sup>, 1.28 &gt;</li> </ul>	XEQ " <b>DLN</b> "	-> D <sup>3.14</sup> Ln (1.28) = 1.138569850
• Exponential: 3.14 ENTER <sup>^</sup> , 1.28	XEQ " <b>DEXP</b> "	-> D <sup>3.14</sup> Exp (1.28) = 3.501963669

A few Special Functions.
• Sine Integral $D^{\mu}$ Si $x = 2^{\mu-2} \pi x^{1-\mu} F_3^{-} (1/2, 1; 3/2, (2-\mu)/2, (3-\mu)/2; -x^2/4)$
• <u>Hyperbolic Sine Integral</u> $D^{\mu}$ Shi $x = 2^{\mu-2} \pi x^{1-\mu} {}_2F^{\sim}_3 (1/2, 1; 3/2, (2-\mu)/2, (3-\mu)/2; x^2/4)$
• <u>Cosine Integral</u> $D^{\mu} Ci x = [FC^{(\mu)}_{log}(x) + \gamma / \Gamma(1-\mu)] x^{-\mu} - 2^{\mu-3} sqrt(\pi) x^{2-\mu} {}_{2}F^{\sim}_{3}(1, 1; 2, (3-\mu)/2, (4-\mu)/2; -x^{2}/4)$
• <u>Hyperbolic Cosine Integral</u> D <sup><math>\mu</math></sup> Chi x = [ FC <sup>(<math>\mu</math>)</sup> <sub>log</sub> (x) + $\gamma$ / $\Gamma$ (1- $\mu$ ) ] x <sup>-<math>\mu</math></sup> + 2 <sup><math>\mu</math>-3</sup> sqrt( $\pi$ ) x <sup>2-<math>\mu</math></sup> <sub>2</sub> F <sup>-</sup> <sub>3</sub> (1, 1 ; 2, (3- $\mu$ )/2, (4- $\mu$ )/2 ; x <sup>2</sup> /4)
• Exponential Integral $D^{\mu}$ Ei x = [ $FC^{(\mu)}_{log}(x) + \gamma / \Gamma(1-\mu)$ ] x $^{-\mu} + x^{1-\mu} _{2}F^{\sim}_{2}(1, 1; 2, 2-\mu; x)$
• <u>Fresnel Cosine Integral</u> $D^{\mu} C(x) = 2^{2\mu \cdot 3/2} \pi^{3/2} x^{1 \cdot \mu}{}_3F^{-}{}_4 [ 1/4 , 3/4 , 1 ; (2 \cdot \mu)/4 , (3 \cdot \mu)/4 , (4 \cdot \mu)/4 , (5 \cdot \mu)/4 ; -\pi^2 x^4/16 ]$
• <u>Fresnel Sine Integral</u> $D^{\mu} S(x) = 2^{2\mu - 11/2} \pi^{5/2} x^{3 - \mu} {}_{3}F^{-}_{4} [ 3/4 , 1 , 5/4 ; (4 - \mu)/4 , (5 - \mu)/4 , (6 - \mu)/4 ; -(PI)^{2} x^{4}/16 ]$
• Spherical Bessel Function - 1st kind $D^{\mu} j_n (x) = 2^{\mu - 2n - 1} \pi x^{n - \mu} \Gamma(n+1) {}_2F^{\sim}_3 [(n+1)/2, (n+2)/2; (n+1-\mu)/2, (n+2-\mu)/2, n+3/2; -x^2/4]$
• Modified Bessel Function - 1st kind $D^{\mu} I_n (x) = 2^{\mu-2n} \operatorname{sqrt}(\pi) x^{n-\mu} \Gamma(n+1) {}_2F^{\sim}{}_3 [(n+1)/2, (n+2)/2; (n+1-\mu)/2, (n+2-\mu)/2, n+1; x^2/4]$
• Bessel Function - 1st kind $D^{\mu} J_n (x) = 2^{\mu-2n} \operatorname{sqrt}(\pi) x^{n-\mu} \Gamma(n+1) {}_2F_3^{-} [(n+1)/2, (n+2)/2; (n+1-\mu)/2, (n+2-\mu)/2, n+1; -x^2/4]$
• <u>Modified Bessel Function - 2nd kind</u> ; where n is not an integer. $D^{\mu} K_{n} (x) = 2^{\mu-2n-1} \pi^{3/2} x^{-\mu-n} \csc(n.\pi) \{ 16^{n} \Gamma(1-n) \ _{2}F^{\sim}_{3} [(1-n)/2, (2-n)/2; (1-\mu-n)/2, (2-\mu-n)/2, 1-n; x^{2}/4] - x^{2n} \Gamma/(n+1) \ _{2}F^{\sim}_{3} [(n+1)/2, (n+2)/2; (n+1-\mu)/2, (n+2-\mu)/2, n+1; x^{2}/4] \}$
• <u>Bessel Function - 2nd kind</u> ; where n is not an integer. $D^{\mu} Y_{n} (x) = 2^{\mu-2n} (PI)^{1/2} x^{-\mu-n} \csc(n.\pi) \{ -16^{n}\Gamma(1-n) {}_{2}F^{\sim}{}_{3} [(1-n)/2, (2-n)/2; (1-\mu-n)/2, (2-\mu-n)/2, (1-n)/2, (1-n)/2, (1-\mu-n)/2, $
• <u>Generalized Laguerre's Functions</u> $D^{\mu} L^{a}_{n}(x) = [\Gamma(n+a+1)/\Gamma(n+1)] x^{-\mu} F^{2}_{2}(1, -n; a+1, 1-\mu; x)$
• <u>Airy Functions</u> $D^{\mu} Ai(x) = 3^{\mu-4/3} x^{-\mu} \{ 3^{2/3} \Gamma(1/3) {}_{2}F^{*}_{3} [ 1/3 , 1 ; (1-\mu)/3 , (2-\mu)/3 , (3-\mu)/3 ; x^{3}/9 ] - x \Gamma(2/3) {}_{2}F^{*}_{3} [ 2/3 , 1 ; (4-\mu)/3 , (2-\mu)/3 , (3-\mu)/3 ; x^{3}/9 ] \}$
$ \begin{split} D^{\mu} & Bi(x) = 3^{\mu-5/6} x^{-\mu} \left\{ \begin{array}{l} 3^{2/3} \Gamma(1/3) \ _2F^{\sim}_3 \left[ \begin{array}{l} 1/3 \ , 1 \ ; \ (1-\mu)/3 \ , \ (2-\mu)/3 \ , \ (3-\mu)/3 \ ; \ x^{3}/9 \ \right] + x \ \Gamma(2/3) \ _2F^{\sim}_3 \left[ \begin{array}{l} 2/3 \ , 1 \ ; \ (4-\mu)/3 \ , \ (2-\mu)/3 \ , \ (3-\mu)/3 \ ; \ x^{3}/9 \ \right] \right\} \end{split} $
• Error Function $D^{\mu} \operatorname{Erf} (x) = 2^{\mu} x^{1-\mu} {}_2F^{\sim}_2 [ 1/2 , 1 ; (2-\mu)/2 , (3-\mu)/2 ; -x^2 ]$

 $\begin{array}{l} \bullet \quad \underline{\text{Hermite Function}} \\ D^{\mu} \; H_n \; (x) = \left[ \begin{array}{c} 2^{n+\mu} \; (\pi) \; x^{-\mu} \; / \; \Gamma((1\!-\!n)/2) \end{array} \right] \, _2F^{\sim}_2 \left[ \begin{array}{c} 1 \; , \; -n/2 \; ; \; (1\!-\!\mu)/2 \; , \; (2\!-\!\mu)/2 \; ; \; x^2 \; \right] - \left[ \begin{array}{c} 2^{n+\mu} \pi \; x^{1-\mu} \; / \; \Gamma(((-n)/2) \; \right] \, _2F^{\sim}_2 \left[ \begin{array}{c} 1 \; , \; (1\!-\!n)/2 \; ; \; (1\!-\!\mu)/2 \; ; \; (3\!-\!\mu)/2 \; ; \; x^2 \; \right] \end{array} \right]$ 

• <u>Kummer's Function</u>  $D^{\mu} F(a;b;x) = x^{-\mu} \Gamma(b) {}_{2}F^{-}_{2}(1, a; 1-\mu, b; x)$ 

#### **Examples:**

• Sine Integral 3.14, ENTER<sup>^</sup>, 1.28, XEQ "DSI"  $\rightarrow D^{3.14}$  Si (1.28) = -0.045395644 • Hyperbolic Sine Integral 3.14, ENTER<sup>^</sup>, 1.28, XEQ "DSHI"  $\rightarrow D^{3,14}$  Shi (1.28) = 0.576495211 Cosine Integral 3.14, ENTER<sup>^</sup>, 1.28, XEQ "DCI"  $\rightarrow D^{3.14}$  Ci (1.28) = 1.367323895 Hyperbolic Cosine Integral 3.14, ENTER^, 1.28, XEQ "DCHI" -> D<sup>3.14</sup> Chi (1.28) = 1.405640394 • Exponential Integral  $\rightarrow D^{3.14}$  Ei (1.28) = 1.982135606 3.14, ENTER^, 1.28, XEQ "DEI" • Fresnel Cosine Integral 3.14, ENTER^, 1.28, XEQ "DCX" -> D<sup>3.14</sup> C (1.28) = 16.95612253 • Fresnel Sine Integral 3.14, ENTER<sup>^</sup>, 1.28, XEQ "DSX"  $\rightarrow D^{3.14}$  S (1.28) = -11.20302776 • Spherical Bessel Function - 1st kind 3.14, ENTER<sup>^</sup>, 2.41, ENTER<sup>^</sup>, 1.28 XEQ "DSB1" ->  $D^{3.14} i_{2.41} (1.28) = -0.064451622$ • Modified Bessel Function - 1st kind , n # -1 , -2 , -3 , ... 3.14, ENTER<sup>^</sup>, 2.41, ENTER<sup>^</sup>, 1.28, XEQ "DINX" ->  $D^{3.14}$  I<sub>2.41</sub> (1.28) = 0.352247279 • Bessel Function - 1st kind , n # -1 , -2 , -3 , ... 3.14, ENTER<sup>^</sup>, 2.41, ENTER<sup>^</sup>, 1.28 XEQ "DJNX" ->  $D^{3,14} J_{2,41}$  (1.28) = -0.150524582 • Modified Bessel Function - 2nd kind - non-integer order 3.14, ENTER<sup>^</sup>, 2.41, ENTER<sup>^</sup>, 1.28 XEQ "DKNX" ->  $D^{3.14}$  K<sub>2.41</sub> (1.28) = -38.98469314 • Bessel Function - 2nd kind - non-integer order 3.14, ENTER<sup>^</sup>, 2.41, ENTER<sup>^</sup>, 1.28, XEQ "DYNX" ->  $D^{3.14} Y_{2.41} (1.28) = 25.49308580$ • Generalized Laguerre's Functions 3.14, ENTER^, 1.76, ENTER^, 2.41, ENTER^, 1.28 XEQ "DLANX" ->  $D^{3.14} L^{1.76}_{2.41}$  (1.28) = -1.767203465

• Airy Functions 3.14 ENTER^, 1.28 XEQ "DAIRY" ->  $D^{3.14}$  Ai (1.28) = -0.162004857 X<>Y  $D^{3.14}$  Bi (1.28) = 3.432592624

- Error Function 3.14 ENTER^, 1.28, XEQ "DERF" ->  $D^{3.14}$  Erf ( 1.28 ) = 1.250557023
- Hermite Function 3.14 ENTER^, 2.41, ENTER^, 1.28 XEQ "DHMT" ->  $D^{3.14} H_{2.41}$  ( 1.28 ) = 3.537707646
- Kummer's Functions With a = sqrt(2) & b = sqrt(3) 2, SQRT, STO 01, 3, SQRT, STO 02 3.14 ENTER^, 1.28, XEQ "DKUM" ->  $D^{3.14} F(2^{1/2}; 3^{1/2}; 1.28) = 2.075891500$

End of the Manual.-

